

Green Flash

High performance computing for real-time science



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Observing stars from the ground

Atmospheric turbulence reduces image quality





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Observing stars from the ground

• From a spherical wave to a flat wavefront













Observing stars from the ground

- Mixture of hot and cold air disturbs the wavefront
- Building a larger telescope does not help to get sharper images



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CROGATE



207



 $\rightarrow \downarrow \lambda/D$



 Compensate in real-time for the wavefront errors ...





... using one or several deformable mirror(s) ...





MICROGATE





 while wavefront errors are measured using one or several wavefront sensors

Turbulent wavefront





Green Adaptive optics

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Quick demo using end-to-end simulation tool





 Astronomers revenge: full resolution of a 8m telescope

Moon surface at $\lambda = 2.3$ microns (NAOS)



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- Ċ Image blurred by 2015
 - turbulence

26" = 45 km



Green European Extremely Large Telescope

39m diameter telescope : x5 in diameter
=> x25 in system complexity

- 100m dome, 2800 tones structure rotating @ 360°, seismic safe (Chile)
- 1.2 G€ project, first light foreseen in 2024
- Construction led by ESO (European Southern Observatory), international organization funded by 15 European countries
- Telescope components + science instruments built by European research labs + industrial partners







Green Addressing grand challenges

• First stars, first galaxies, birth and evolution of the Universe



GOODS South Field • WFC3 Early Release Science Data Hubble Space Telescope • WFC3/UVIS/IR • ACS/WFC

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NASA, ESA, R. Windhorst (Arizona State University), P. McCarthy (Carnegie Institution of Washington), R. O'Connell (University of Virginia), and the WFC3 Science Oversight Committee

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STScI-PRC10-01a

• From classical AO ...







• From classical AO ...

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• From classical AO ... to multi-conjugate AO (MCAO) ...











• WFS needs a bright star ... when not available => create your own !











From Laser guided AO... to Laser-tomography AO ... •

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• multiple guide stars and multiple WF correctors

(c) Observing the GOODS South cosmological field with MOAO.

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Green Adaptive optics flavors

... to ground-layer AO ... to multi-object AO (MOAO) !

Ground layer AO (GLAO) One DM in telescope pupil

Multiobject AO (MOAO) one DM per observed object

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Green E-ELT: an adaptive telescope

... with multiple Laser guide stars ...

Green E-ELT: an adaptive telescope

- ... and multiple technical challenges !
 - Telescope components (dome, mirror segments, etc ...)
 - Fast and low noise detectors
 - for science channel (in the IR domain) but also for AO (in the visible domain, for WFS)
 - New optical components
 - science instrument but also AO WFS
 - Compute technologies
 - for the **real-time control** (RTC) of the telescope and AO instruments
 - for the **simulation** (instrument design trade-off studies)
- Limited budget for overall instruments set : ~100M€ including staff effort for operations

Green AO RTC concept

Green AO RTC concept : data pipeline

AO RTC concept : smart interconnect

Green AO RTC concept : supervisor

Introduction to Green Flash

- Program objectives: 3 research axes
 - 2 technological developments and 1 validation study
- Real-time HPC using accelerators and smart interconnects
 - Assess the determinism of accelerators performance
 - Develop a smart interconnect strategy to cope for strong data transfer bandwidth constraints
- Energy efficient platform based on FPGA for HPC
 - Prototype a main board, based on FPGA SoC and PCIe Gen3
 - Cluster such boards and assess performance in terms of energy efficiency and determinism
- AO RTC prototyping and performance assessment
 - Assemble a full functionality prototype for a scalable AO RTC targeting the MAORY system
 - Compare off-the-shelf solutions based on accelerators and new FPGA-based concept

01/22/2016

Assessing new HPC concepts

Green Flash project

- EU funded Partners
 - 2 academic partners
 - LESIA, Observatoire de Paris, P.I. Damien G.
 - CfAI, University of Durham
 - 2 industrial partners
 - Microgate : Italian SME designing FPGA solutions for various applications (including astronomical AO)
 - PLDA: French SME developing FPGA solutions (mostly IP cores, world leader in PCIe IPs
- KAUST ECRC as external partner

Green AO RTC concept : data pipeline

FPGA solution : µXcomp

Based on ARRIA 10AX115:

- 1518 DSP blocks
- 6.6MB int. RAM
- 96 XCVR

Board features:

- Optimized for heavy deterministic computation in floating-point
- Large Bandwidth between HMC and FPGA - 4 links 16 lanes/link up to 15Gbps/lane = 120GB/s bidirectional
- Extremely low jitter
- More power efficient compared to GPUs
- Offers a lot of different interfaces on board or via the FMC connector and extension cards

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FPGA solution : μ Xlink

ARRIA 10AS066 SoC:

- 1.5GHz ARM dual-core Cortex-A9 on-chip processor
- 1855 DSP blocks
- 5.2MB int. RAM
- max. 48 XCVR

Board features:

- ARM embedded processor for stand-alone real-time box
- Powerful PCle root port because of ARM and OS
- Management of accelerator cards on the PCIe interface
- Running control software using a full OS (e.g. Linux)
- Easy implementation of different communication protocols
- Offers a lot of different interfaces on board or via the FMC connector and extension cards

FPGA solutions: status

The first prototype of the two FPGA boards, the μ XComp board is manufactured and is currently under test. After the validation of the interfaces and the communication between FPGA and HMC some more boards of this type will be produced and made available to the team.

Platform based on accelerators

One generic node architecture, two applications :

- Real-time memory bound linear algebra (AO linear control, a.k.a. real-time pipeline)
- High throughput compute bound linear algebra (AO supervisory tasks, a.k.a. supervisor)

For each application, nodes are interconnected into a cluster. For the full featured prototype, the two clusters are interconnected

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RT data pipeline with GPUs

Prototype using latest generation GPU cluster

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5104

Concept studied at LESIA

Observatoire - LESIA

System dimensioning

MCAO @ E-ELT scale

- POLC control scheme + LGS WFS : 2.5 TMAC/s with 250 Gb/s of streaming data
- Upper limit from instruments specification capture during PDR (actual first light instruments may require less)

	K20C	K40	K80	P100	
B _{theo}	208	288	240 (x2)	732	
B _{no ECC}	175 (84%)	236 (82%)	200 (x2, 83%)	460 (62%)	
B _{ECC}	150 (72%)	208 (72%)	173 (x2, 72%)	460 (62%)	

Memory bandwidth

Number of GPUs required

ECC	K20C	K40	K80	P100
Off	12	9	6	5
On	14	10	6	5

Multi-GPU prototype

Persistent kernel implementation

Persistent kernel implementation

Strong scalability

Constant case with 10,048 slopes x 15,000 commands

Histogram

Case with 10,048 slopes x 15,000 commands on 4 devices

Average : 0.45ms Jitter : 17µs

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Data acquisition

FPGA writes/reads directly to/from GPU memory Using only writes would be better though

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FPGA/GPU optimized sync.

Little to no improvements, but CPU free for other kind of computations

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Data acquisition + persistent kernels

FPGA PLDA XPressG5 GPU Tesla C2070 OS Debian wheezy

Camera EVT HS-2000M 10GbE network

Green AO RTC concept : supervisor

Mix of cost function optimization for parameters identification ("Learn" process) and linear algebra for reconstructor matrix computation ("apply" process)

l'Observatoire LESIA

KAUST

Parameters identification ("Learn" process) 200

- Fitting measurements covariance matrixon on a model including system and turbulence parameters
- Using a score function

$$F(x) = \sum_{k=1}^{N^2} [Cmm_k - f_k(x)]^2$$

- Levenberg-Marquardt algorithm for function optimization
- Exemple of turbulence profile reconstruction

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• Dual stage process (5 layers + 40 layer

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Performance for parameters identification ("Learn" process) Multi-GPU process, including matrix generation and LM fit Time to solution for a matrix size of 86k :240s (4 minutes)

- first pass (5 layers) : 25s
- Second pass (40 layers) : 213s

Performance for parameters identification ("Learn" process) Multi-GPU process, including matrix generation and LM fit Time to solution for a matrix size of 86k :

- first pass (5 layers) : 25sec
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Reconstructor matrix computation ("apply" process)

• Compute the tomographic reconstructor matrix using covarince matrix between "truth" sensor and other WFS and invert of measurements covariance matrix

 $R' = Ctm \cdot Cmm_f^{-1}$

- Can use various methods. "Brute" force : direct solver
- Standard Lapack routine : "posv" : mostly compute-bound, high level of scalability
- Highly portable code : explore various architectures by using standard vendor provided maths libraries

Performance for reconstructor matrix computation ("apply" process)

Comparing last generation of GPU (NVIDIA P100) and last generation of Intel Xeon Phi (KNL)

8 GPUs together reach more than 21 TFLOP/s while a single KNL can only reach about 1.2 TFLOP/s in peak performance

Performance for reconstructor matrix computation ("apply" process)

Comparing last generation of GPU (NVIDIA P100) and last generation of Intel Xeon Phi (KNL)

GPUs can deliver better peak perf. (saturation not reached, expect >2.5 or more) and the NVIink interconnect seems to perform very well

Performance for reconstructor matrix computation ("apply" process)

 Comparing last generation of GPU (NVIDIA P100) and last generation of Intel Xeon Phi (KNL)

 Record time-to-solution on DGX-1 : MAORY / HARMONI full scale (100k x 100k matrix) : 25sec to compute tomographic reconstructor

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Performance evolution of ToR computation over the past years

Critical for AO instrument design studies

Goal is to produce an observation forecast •

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Green AO RTC concept : data streams

Green AO RTC concept : local / global interco.

Smart interconnect architecture

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Smart interconnect concept

Eased devel.
process
using the
QuickPlay
tool from
PLDA

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QuickPlay

QuickPlay[™]

Introducing QuickPlay

QuickPlay^{**}

FPGA Design with QuickPlay IDE

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- MODEL
 - C/C++ functional modeling

VERIFY & VALIDATE

Desktop execution of system functional model

BUILD

Hardware implementation: HLS, Logic Synthesis, P&R

EXECUTE

FPGA based system hardware execution

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QuickPlay Hardware Accelerator Abstraction Layer

Smart interconnect prototyping

- Single generic design / multiple target boards
 - ExpressK-US board (hosting a Kintex UltraScale from Xilinx)
 - ExpressGX V board (hosting a Stratix V from Altera)
 - μXlink board from microgate (hosting a Arria 10 board from Altera)

Smart interconnect prototyping

International joint academic-industrial project funded by EU

- Key contributor to European Extremely Large Telescope project
- Exploiting emerging computing technologies (both HW and SW)
- KAUST-ECRC as external partner
- Energy efficiency plays a critical role (remote location, limited operational budget)
 - Accelerator-based system architecture
 - Optimized data streaming using FPGA technology

Optimized Linear algebra is key to the various sub-systems

- Not only for system operations but also simulations for the system design
- Original approach designed in collaboration with the team at KAUST

Already enhancing the readiness level of commercial solutions

- Contribution to QuickPlay development environment from PLDA
- Design of innovative FPGA boards at Microgate

